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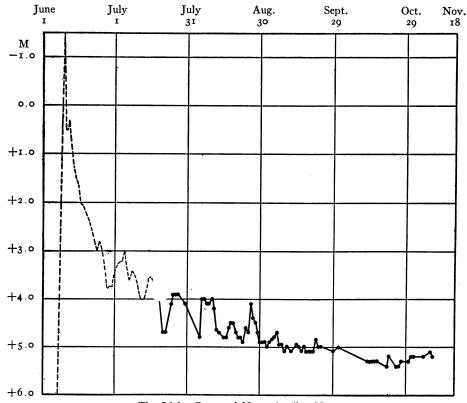
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This parallax gives for the absolute magnitude of the star, +14.3 photovisual and +14.8 photographic. It is, therefore, by far the faintest F-type star known at the present time.

A. VAN MAANEN.

## THE LIGHT CURVE OF NOVA AQUILAE No. 3

The following observations of the magnitude of *Nova Aquilae No. 3* were made visually on Mount Wilson, using the comparison stars listed in *Harvard Bulletin*, No. 661. Most of the estimates were made when the nova was near the zenith, and no correction has been applied for atmospheric extinction. In the diagram the light curve for dates earlier than July 17 is taken from *Harvard Circular*, No. 208. The observations on July 21 and July 22, which show a deep minimum, were confirmed by the estimates of other observers on Mount Wilson. On August 12,  $\beta$  *Scuti* was recorded as appearing brighter than magnitude 4.5.



The Light Curve of Nova Aquilae No. 3

Date	G.M.T.	Mag.	Remarks	Dat	e	G.N	И.Т.	Mag.	Remarks
Date  July 17 18 19 21 22 24 25 26 27 30 Aug. 5 6 7 8 9 10 11 12 13 15 16 17 18 19 20 21 22 23 24 25 26 27 28	G.M.T.  16h 55m 17 0 17 20 16 40 17 15 16 40 17 15 16 40 17 30 16 30 16 55 17 10 16 55 17 0 16 55 17 0 16 55 17 0 16 55 17 0 16 55 17 0 16 50 16 50 16 50 16 50 16 50 16 50 17 35 16 17 0 17 10 18 15 45 19 10 19 10	3.78 3.97 4.71 3.99 3.91 4.80 4.11 4.02 55 4.78 4.84 4.55 4.78 4.88 4.96 4.47	Moonlight  " " " " " " " " " " " " " " " " " "	Sept.		16 <sup>h</sup> 16 16 16 16 16 16 17	15 <sup>m</sup> 0 15 10 0 20 20 0 40 40 0	4.85 4.85 4.95 4.95 5.01 5.01 5.01 5.01 5.01 5.01 5.01 5.0	Some cirrus Thru clouds Moonlight  " ";scattered clouds Thru clouds Scattered clouds Moonlight  " " Thru clouds
29 30 31 Sept. 1	16 55 17 30 16 40 17 10 16 50	4.7 4.9 4.9 4.9 5.0	Scattered clouds	Nov.	31 5 8 9	16 16 15 15	5 10 0 10 55	5.2 5.2 5.2 5.1 5.2	

MILTON HUMASON.

## THE MOTION IN SOME A DOUBLE STARS

## FOURTH NOTE

Three earlier notes<sup>1</sup> in these Publications have called attention to certain binary stars discovered in the Lick Observatory double-star survey which were shown to be in relatively rapid orbital

<sup>&</sup>lt;sup>1</sup>Publ. A. S. P., **27**, 230, 1915; **28**, 276, 1916; **30**, 69, 1918.